

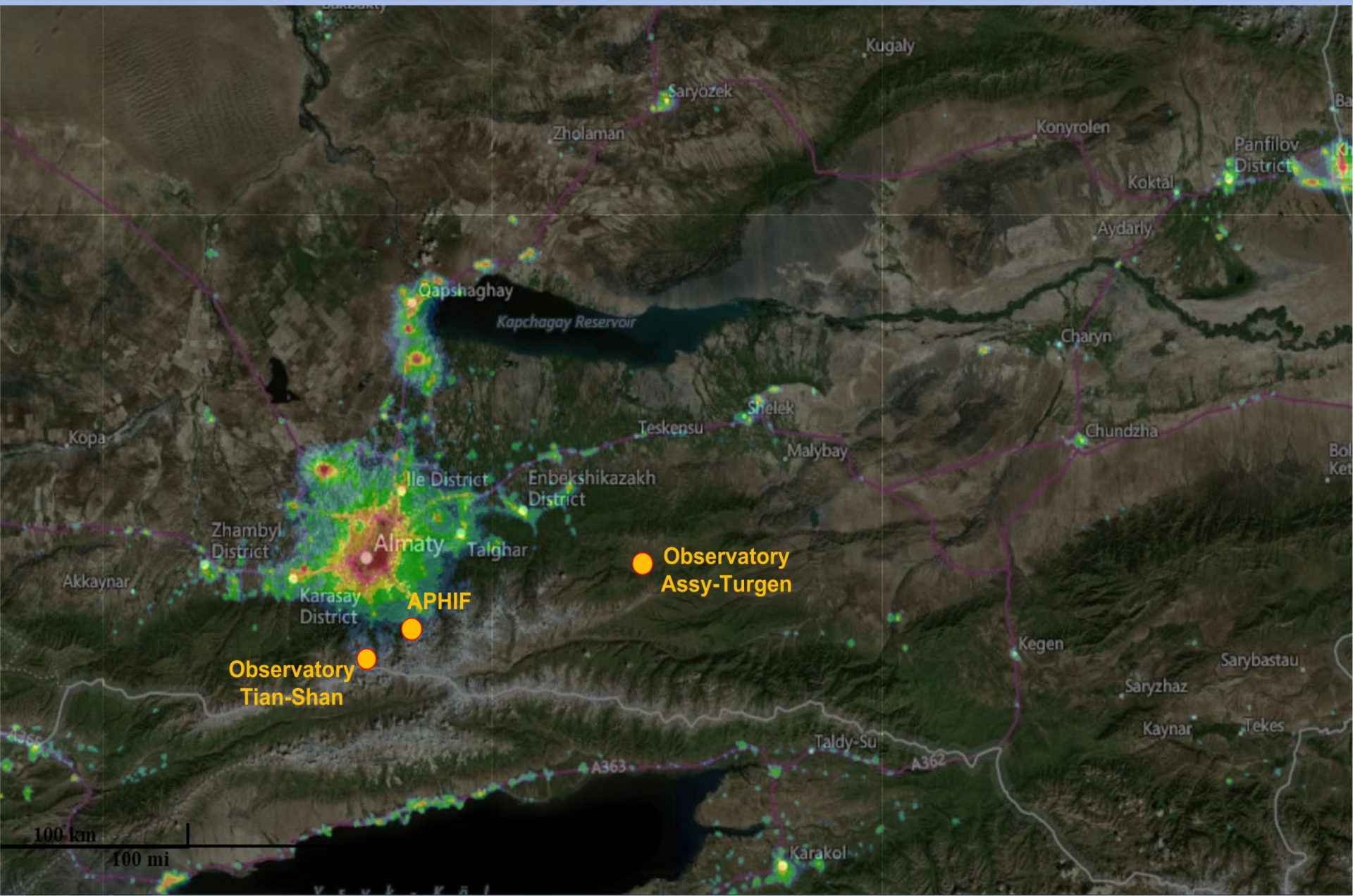


Practical use and capabilities of the AZT-20 telescope of Assy-Turgen observatory

Kim V.Yu ^{1,2}

1. Higher school of economics (Moscow, Russia) (postdoc)
2. Fesenkov astrophysical institute (Almaty, Kazakhstan)

Observatories APhi



Assy-Turgen Astronomical Observatory



1,5m telescope AZT-20

1,0m telescope Zeiss 1000

<https://goo.gl/maps/C0nKW>

2700 meters above sea level

Mounting of AZT-20 dome



Astroclimate of Assy-Turgen Plateau

- Median astronomical seeing observed during 1976-1981 was 0.64'' (minimum 0.20'')
- Clear night: 178 per year (1500 hours)
- Wind speed in clear nights is low: 1,7 m/sec
- Night median temperature:
summer + 9.5°C
winter – 10.5°C
- Horizon openness: 3,5°
- Minimal levels of sky brightness of V mag/arcsec²:

V mag/arcsec ² :	Photometry system
22,18	B
21,34	V
20,42	R

- UV atmospheric transmission

λnm	301	313,5	345	412
P	0,082	0,340	0,575	0,735

AZT-20 1,5m telescope

D = 1560 mm

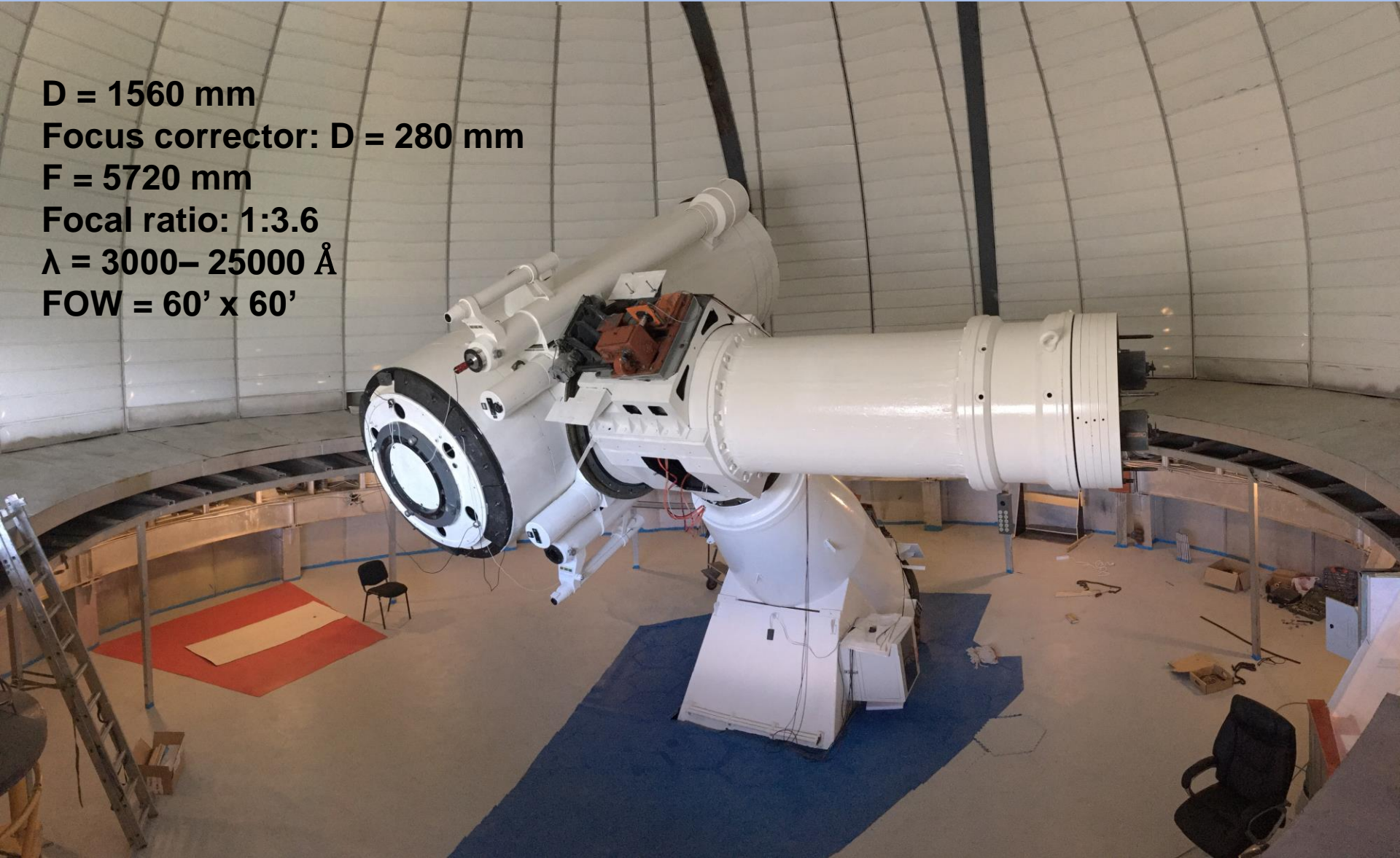
Focus corrector: D = 280 mm

F = 5720 mm

Focal ratio: 1:3.6

$\lambda = 3000\text{--}25000 \text{ \AA}$

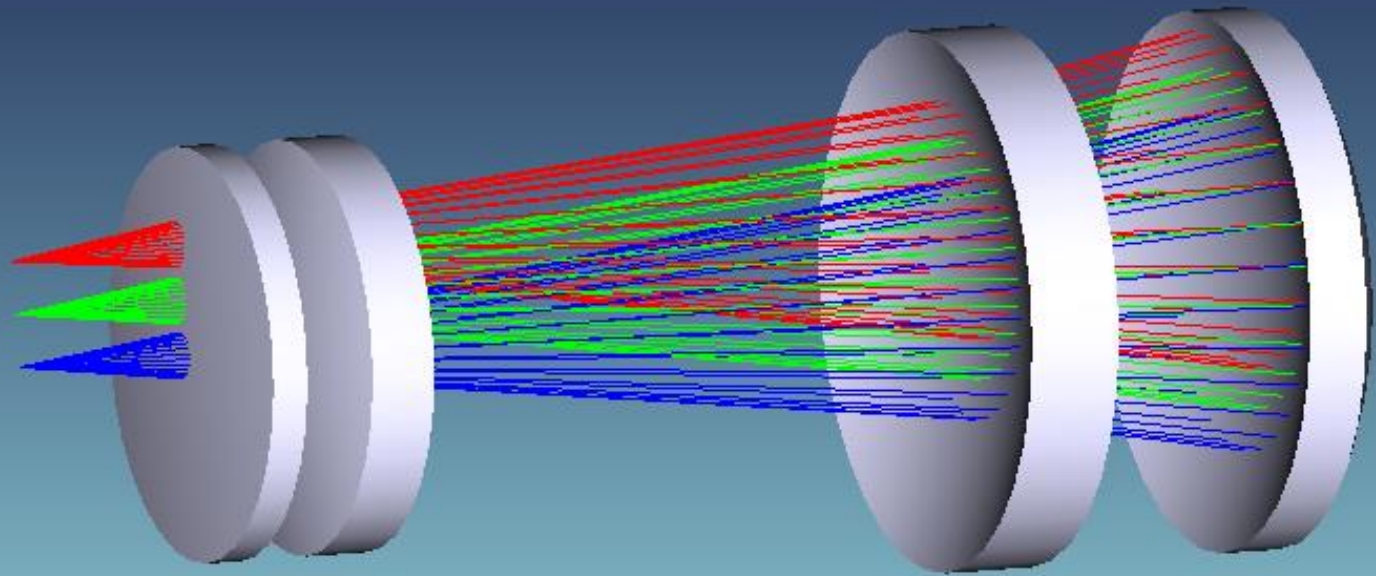
FOV = 60' x 60'

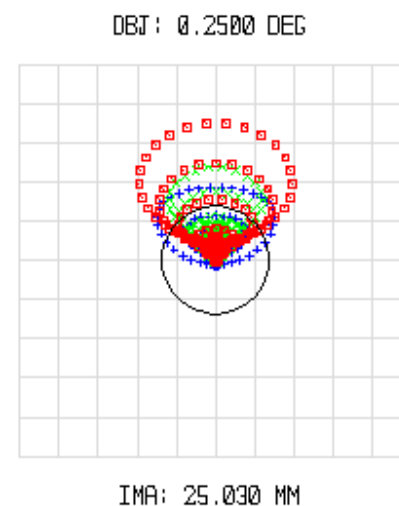
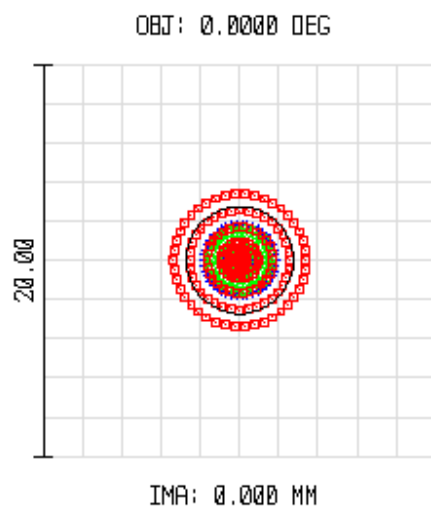


AZT-20 1,5m telescope

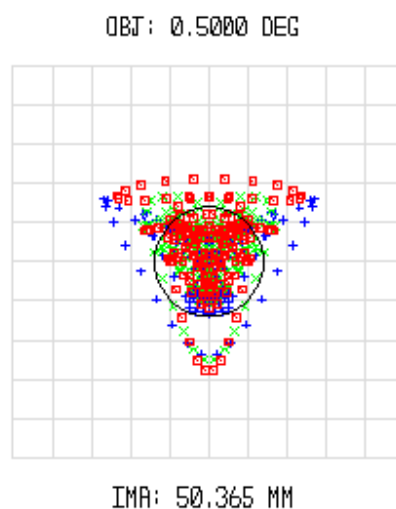
3: Shaded Model Z = -451.65854, Y = -193.03178

Update Settings Print Window Text Zoom Spin





+ 0.4861
× 0.5876
◻ 0.6563



SURFACE: IMA

SPOT DIAGRAM

WED OCT 18 2017 UNITS ARE μm , AIRY RADIUS : 2.736 μm

FIELD : 1 2 3

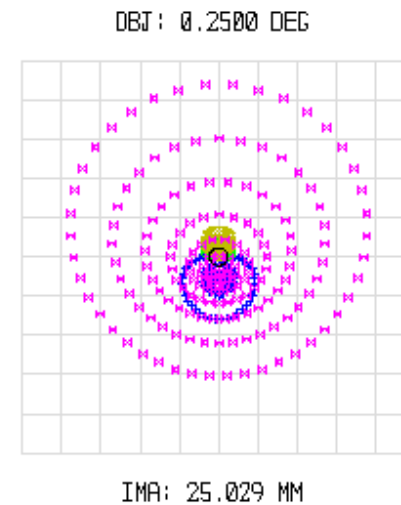
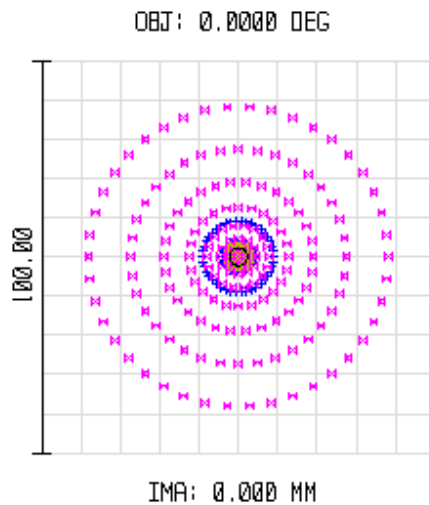
RMS RADIUS : 1.539 2.164 2.507

GEO RADIUS : 3.373 7.004 6.220

SCALE BAR : 20

REFERENCE : CHIEF RAY

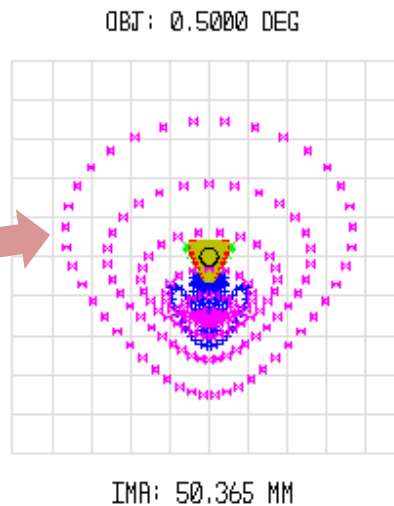
AZT20_WYNNE_310817_INTANE.ZMX
CONFIGURATION 1 OF 1



+ 0.3100
 × 0.4876
 ◻ 0.5563
 ✕ 0.6500
 H 2.5000

UBVRIJK

As far IR
 pixels bigger
 size than UV-
 Vis tech



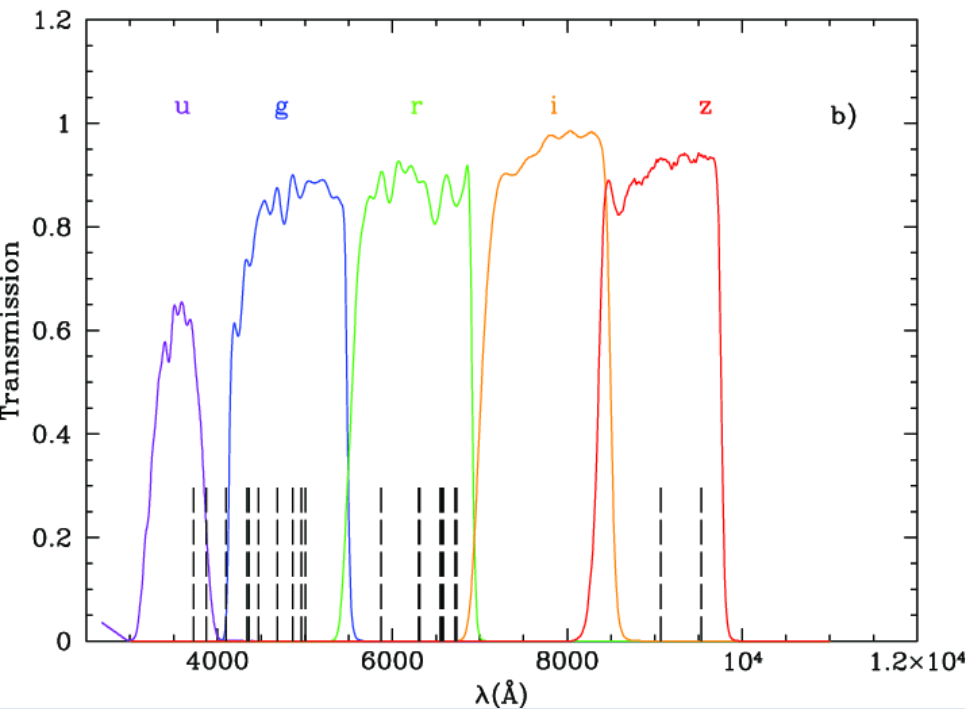
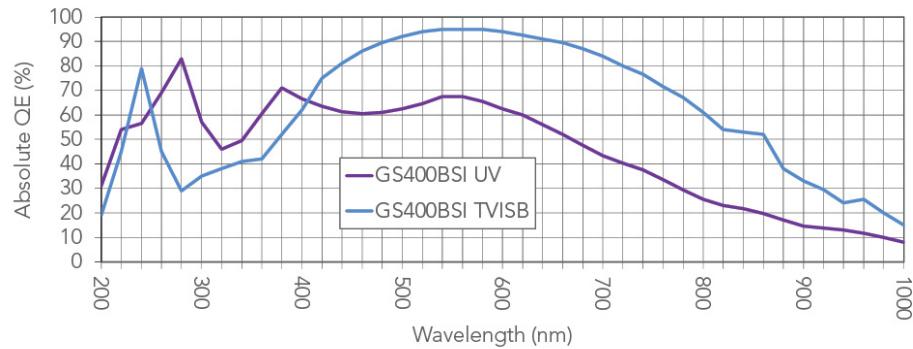
SURFACE: IMA

SPOT DIAGRAM

WED OCT 18 2017 UNITS ARE μm , AIRY RADIUS : 2.27 μm
 FIELD : 1 2 3
 RMS RADIUS : 11.394 11.474 12.113
 GEO RADIUS : 38.172 43.811 38.137
 SCALE BAR : 100 REFERENCE : CHIEF RAY

AZT20_WYNNE_310817_INTANE.ZMX
 CONFIGURATION 1 OF 1

Equipment for photometry



CMOS back illuminated camera
Kepler KL400

Tasks for photometric observations on AZT-20

- 1. Optical afterglow of gamma-ray bursts
 - (published more than 30 telegrams in GCN Circular). Collaboration with SRI RAS
- 2. Wolf-Rayet stars
- 3. AGNs
- 4. etc. (asteroids, satellites)

M100 Galaxy



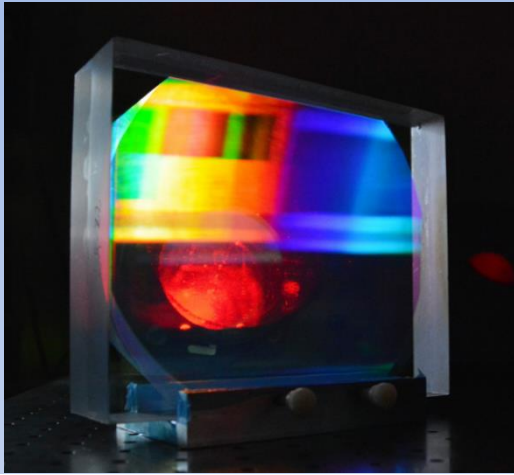
Eagle Nebula



'Hidden' galaxy



Slit Spectrograph for AZT-20 telescope



Volume Phase Holographic (VPH) Gratings

Wasatch) Photonics

Excellent 1st order diffraction efficiency

Broad operational bandwidth with angle tuning

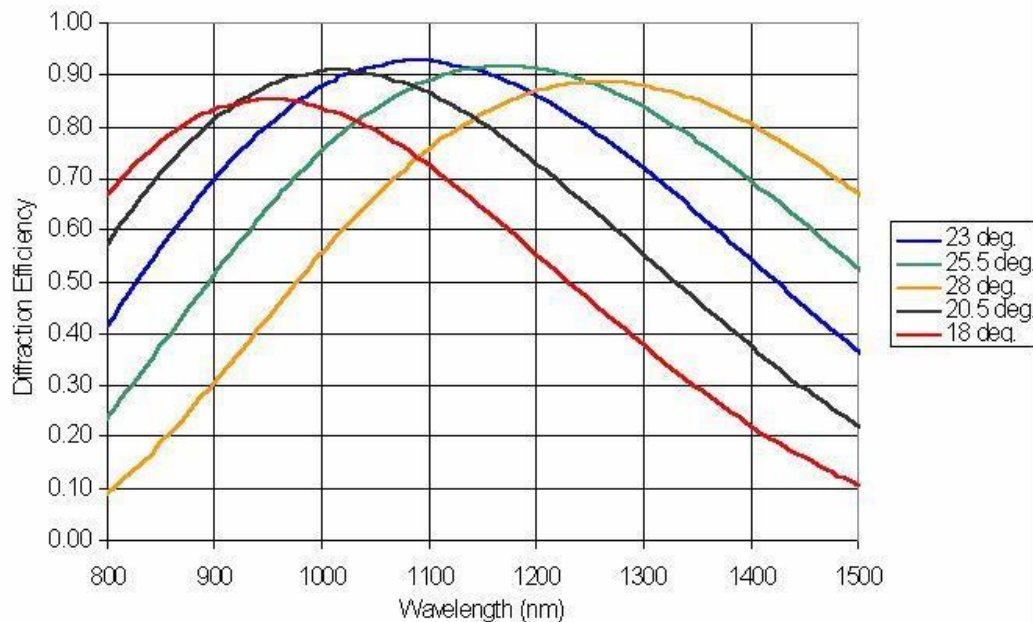
Customizable in wavelength (350-2500 nm)

Gratings (360 and 2400 lines/mm)

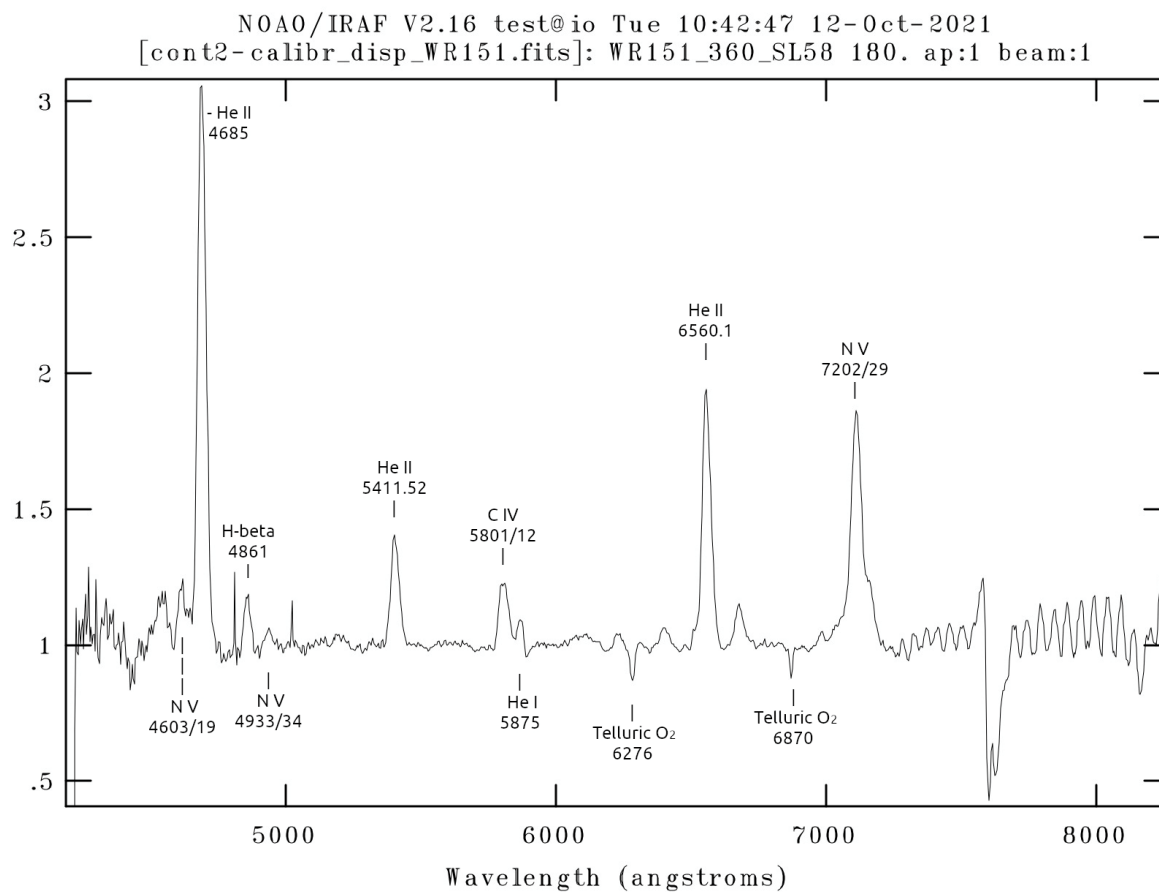
Disp. (4 Å/pix and 0.4 Å/pix)

EmCCD camera Andor (~ 21 mag)

Tunability - Diffraction Efficiency vs Wavelength for Different Angles of Incidence



WR 151 spectrum received on VPH (360) grating



Tasks for spectroscopic observations on AZT-20

- 1. Wolf-Rayet stars
- 2. Gaseous nebulae
- 3. Variable stars (B[e], Ae/Be Herbig's stars, etc)

Further perspectives for AZT-20

- 1. Installation of additional VPH-grating (1700 lines/mm) for H α -region
- 2. Total automation of AZT-20 for observing process.

International collaboration in Assy-Turgen













2021-4-2 13:33



2021-3-31 11:53



2021-3-31 11:53





~~непринятая~~

Система координат: Местная

Топографическая съемка

1:1000

В 1 сантиметре 10 м
Средняя температура воздуха: +10,5 °С
Средняя влажность: 75%

Исполнитель: _____
Объект: _____
Дорожка: _____
Исполнитель: _____
Дата: _____

В 1 сантиметре 10 м
Сплошная горизонтальн. проведена через 0,50 м
Система высот : Балтийская

Экземпль
Объект
Директор
Исполнитель
Дата

Thank you for attention